# CompOSE CompStar Online Supernovae Equations of State

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for the CompOSE core team (Micaela Oertel, Thomas Klähn, S. T.)

**CompOSE** meeting

Institut de Physique Nucléaire de Lyon (IPNL)



# CompOSE I

#### • features

- repository of equations of state (data tables and additional information) for applications in astrophysics, nuclear physics and beyond
- $\circ$  EoS for nuclear/quark matter and stellar matter

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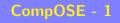
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- information on thermodynamic quantities, chemical composition and microscopic quantities

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- $\circ$  EoS can be one-, two-, or three-dimensional
- information on thermodynamic quantities, chemical composition and microscopic quantities
- $\circ$  flexible data format for storage of EoS tables
- tools for extracting, interpolating and generating EoS tables according to the needs of the user with determination of additional quantities
- supports ASCII and HDF5 data formats in output
- $\circ$  subscription for newsletter available



# CompOSE II

#### • access & up-to-date information

- o website: compose.obspm.fr no registration needed any more
- $\circ$  manual (version 1.00,  $\approx$  70 pages): available from website or arXiv:1307.5715 [astro-ph.SR], will be updated soon



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#### • objective of today's meeting

 $\circ$  extension of CompOSE database  $\Rightarrow$  contribute your EoS



### How to contribute with your EoS

three steps:

#### • preparation of EoS tables

• three tables with parameters (mandatory):

temperature T, baryon number density  $n_b$ , charge fraction  $Y_q$ 

- $\circ$  table with thermodynamic quantities (mandatory)
- table with composition of matter (optional)
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#### • testing and preparation of data sheet

 $\circ$  use of FORTRAN program compose.f90 and C++ program <code>eosform.cpp</code>

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- testing and preparation of data sheet
  - $\circ$  use of FORTRAN program compose.f90 and C++ program <code>eosform.cpp</code>
- uploading your EoS to the database
  - $\circ$  no general rule
  - o contact the CompOSE administrators: develop.compose@obspm.fr

#### • parameter tables (mandatory)

- temperature T
   unit: MeV
   name of file: eos.t
- $\circ$  baryon number density  $n_b$ unit: fm<sup>-3</sup> name of file: eos.nb
- $\circ$  charge fraction  $Y_q$ unit: - (dimensionless) name of file: eos.yq

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- files provide mapping of parameter values  $(X = T, n_b, Y_q)$  to indices  $(i_X)$

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#### • general structure of file:

 $\begin{array}{ll} i_X^{\min} & (\min \min \min index) \\ i_X^{\max} & (\max \min index) \\ X(i_X^{\min}) & (value at \min \min index) \\ X(i_X^{\min}+1) & \\ \vdots & \\ X(i_X^{\max}-1) & \\ X(i_X^{\max}) & (value at \max \min index) \end{array}$ 

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#### • general structure of file:

 $i_X^{\min}$ (minimum index)  $i_X^{\max}$ (maximum index)  $X(i_X^{\min})$ (value at minimum index)  $X(i_X^{\min}+1)$  $X(i_X^{\max}-1)$  $X(i_X^{\max})$ (value at maximum index)  $\circ$  one entry per row,  $X(i_X+1) > X(i_X)$  $\circ i_X^{\max} - i_X^{\min} + 3$  rows in total  $\circ i_X^{\max} - i_X^{\min} + 1$  values of quantity X

- table with thermodynamic quantities (mandatory)
  - $\circ$  name of file: <code>eos.thermo</code>
  - $\circ$  general structure of file

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- o name of file: eos.thermo
- general structure of file

$$m_n \quad m_p \quad I_l$$
 (first row)

$$i_T \ i_{n_b} \ i_{Y_q} \ Q_1 \ Q_2 \ Q_3 \ Q_4 \ Q_5 \ Q_6 \ Q_7 \ N_{\text{add}} \underbrace{q_1 \ q_2 \ \dots \ q_{N_{\text{add}}}}_{N_{\text{add}} \text{ quantities}}$$

(subsequent rows)

#### $\circ$ $m_n$ : neutron mass in MeV, $m_p$ proton mass in MeV

- $\circ I_l$ : lepton index,  $I_l = 0$  no leptons,  $I_l$  with leptons (e and/or  $\mu$ )
- $\circ$  parameter indices:  $i_T i_{n_h} i_{Y_a}$
- $Q_1 = p/n_b$  [MeV],  $Q_2 = s/n_b$  [dimensionless],  $Q_3 = \mu_b/m_n 1$  [dimensionless],  $Q_4 = \mu_q/m_n$  [dimensionless],  $Q_5 = \mu_l/n_b$  [dimensionless],
  - $Q_6 = f/(n_b m_n) 1$  [dimensionless],  $Q_7 = e/(n_b m_n) 1$  [dimensionless]
- $\circ N_{add}$  number of additional quanties  $q_1, \ldots, q_{N_{add}}$  (defined by contributor of EoS) o order of rows 2, 3, . . . irrelevant

#### • table with composition of matter (optional)

- $\circ$  name of file: <code>eos.compo</code>
- general structure of file (every row)

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$$i_{T} i_{n_{b}} i_{Y_{q}} I_{\text{phase}} N_{\text{pairs}} \underbrace{I_{1} Y_{I_{1}} \dots I_{N_{\text{pairs}}} Y_{I_{N_{\text{pairs}}}}}_{N_{\text{pairs}} pairs} X_{\text{pairs}}$$

$$N_{\text{quad}} \underbrace{I_{1} A_{I_{1}}^{\text{av}} Z_{I_{1}}^{\text{av}} Y_{I_{1}} \dots I_{N_{\text{quad}}} A_{I_{N_{\text{quad}}}}^{\text{av}} Z_{I_{N_{\text{quad}}}}^{\text{av}} Y_{I_{N_{\text{quad}}}}}}_{N_{\text{quad}} q_{\text{uadruples}}}$$

• parameter indices:  $i_T i_{n_b} i_{Y_q}$ , phase index  $I_{\text{phase}}$  (defined by contributor of EoS)

- $\circ$  number of pairs  $(N_{
  m pairs})$  and quadruples  $(N_{
  m quad})$ , can change from row to row
- particle (for pairs) or group (for quadruples) indices  $I_i$  (see below) and corresponding number fractions  $Y_i = n_i/n_b$  [dimensionless]
- $\circ$  average mass numbers  $A_{I_i}^{\mathrm{av}}$  and average charge numbers  $Z_{I_i}^{\mathrm{av}}$
- $\circ$  order of rows irrelevant

- table with microscopic quantities (optional)
  - $\circ$  name of file: <code>eos.micro</code>
  - $\circ$  general structure of file (every row)

$$i_T \ i_{n_b} \ i_{Y_q} \ N_{qty} \underbrace{K_1 \ q_{K_1} \ K_2 \ q_{K_2} \ \dots \ K_{N_{qty}} \ q_{K_{N_{qty}}}}_{N_{qty} \text{ pairs}}$$

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- $\circ$  parameter indices:  $i_T \ i_{n_b} \ i_{Y_q}$
- $\circ$  number of quantities  $(N_{
  m qty})$ , can change from row to row
- $\circ$  indices  $K_i$  defining quantity and particle (see below)
- $\circ$  order of rows irrelevant

- indices for identification of particles (here the most common in EoS tables)
  - $\circ$  leptons:  $I_i = 0$  electrons,  $I_i = 1$  muons
  - $\circ$  baryons:  $I_i=10$  neutrons,  $I_i=11$  protons
  - o nuclei: *I<sub>i</sub>* = 1000 *A<sub>i</sub>* + *Z<sub>i</sub>* e.g. *I<sub>i</sub>* = 2001 <sup>2</sup>H, *I<sub>i</sub>* = 3001 <sup>3</sup>H, *I<sub>i</sub>* = 3002 <sup>3</sup>He, *I<sub>i</sub>* = 4002 <sup>4</sup>He
  - $\circ$  more in table 3.2 of manual

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  - $\circ$  nuclei:  $I_i = 1000 \; A_i + Z_i$  e.g.  $I_i = 2001 \; ^2 {\rm H}$ ,  $I_i = 3001 \; ^3 {\rm H}$ ,  $I_i = 3002 \; ^3 {\rm He}$ ,  $I_i = 4002 \; ^4 {\rm He}$
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#### indices for identification of quantities and particles

- $\circ K_i = 1000 I_i + J_i$  with particles index  $I_i$  and quantity index  $J_i$
- $\circ$  quantity indices
  - $J_i = 50$  nonrelativistic single-particle potential  $U_{I_i}$  [MeV]
  - $J_i = 51$  relativistic vector self-energy  $V_{I_i}$  [MeV]
  - $J_i = 52$  relativistic scalar self-energy  $S_{I_i}$  [MeV]

more in table 7.3 of manual

 $\circ$  example:  $K_i = 11052$  scalar self-energy of proton

### Testing of tables and preparation of data sheet

- preparation of FORTRAN program compose.f90
  - $\circ$  download code.zip from CompOSE website
  - o unzip code.zip
  - $\circ$  change line 26 in Makefile to HDF5 = 0
  - $\circ$  compile program with make

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#### • test of EoS table

- o copy your EoS files eos.t, eos.nb, eos.yq, eos.thermo, ... into directory with compose program
- $\circ$  run compose
- maybe you have to modify the sample files eos.parameters and/or eos.quantities (see manual) in case of errors
- $\circ$  a file <code>eos.report</code> should have been generated with information on the EoS

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#### • preparation of data sheet

- $\circ$  download <code>eosform.zip</code> from CompOSE website
- $\circ$  compile C++ program eosform.cpp (e.g. with g++ -o eosform eosform.cpp)
- $\circ$  run <code>eosform</code> to generate <code>LATEXfile datasheet.tex</code>
- $\circ$  compile datasheet.tex and edit if necessary

# Interpolation

- should reproduce values of all quantities at basic grid points
- depends on order I (set in file eos.parameters)
  - $\circ~I=1$  interpolation continuous in function values
  - $\circ~I=2$  interpolation continuous in function and first derivatives
  - $\circ~I=3$  interpolation continuous in function, first and second derivatives
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- present scheme:
  - $\circ$  first step: one-dimensional interpolation in  $Y_q$
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- details in Appendix A of manual
- problems:
  - $\circ$  thermodynamic consistency, quantities not independent
  - $\circ$  oscillations depending on grid resolution
  - $\Rightarrow$  interpolation errors
- alternatives?